

A Catalogue of Geometrically-Modelled Coronal Mass Ejections Observed by the STEREO Heliospheric Imagers

D. Barnes¹, J. Byrne¹, J. Davies¹, R. Harrison¹, C. Perry¹,
C. Möstl², A. Rouillard³, V. Bothmer⁴, L. Rodriguez⁵, J. Eastwood⁶,
E. Kilpua⁷, P. Gallagher⁸, D. Odstrčil⁹

(1) STFC, Rutherford Appleton Laboratory, UK;

(2) University Graz, Austria;

(3) Paul Sabatier Université, France;

(4) University of Göttingen, Germany;

(5) Royal Observatory of Belgium, Belgium;

(6) Imperial College London, UK;

(7) University of Helsinki, Finland;

(8) Trinity College Dublin, Ireland;

(9) George Mason University, USA

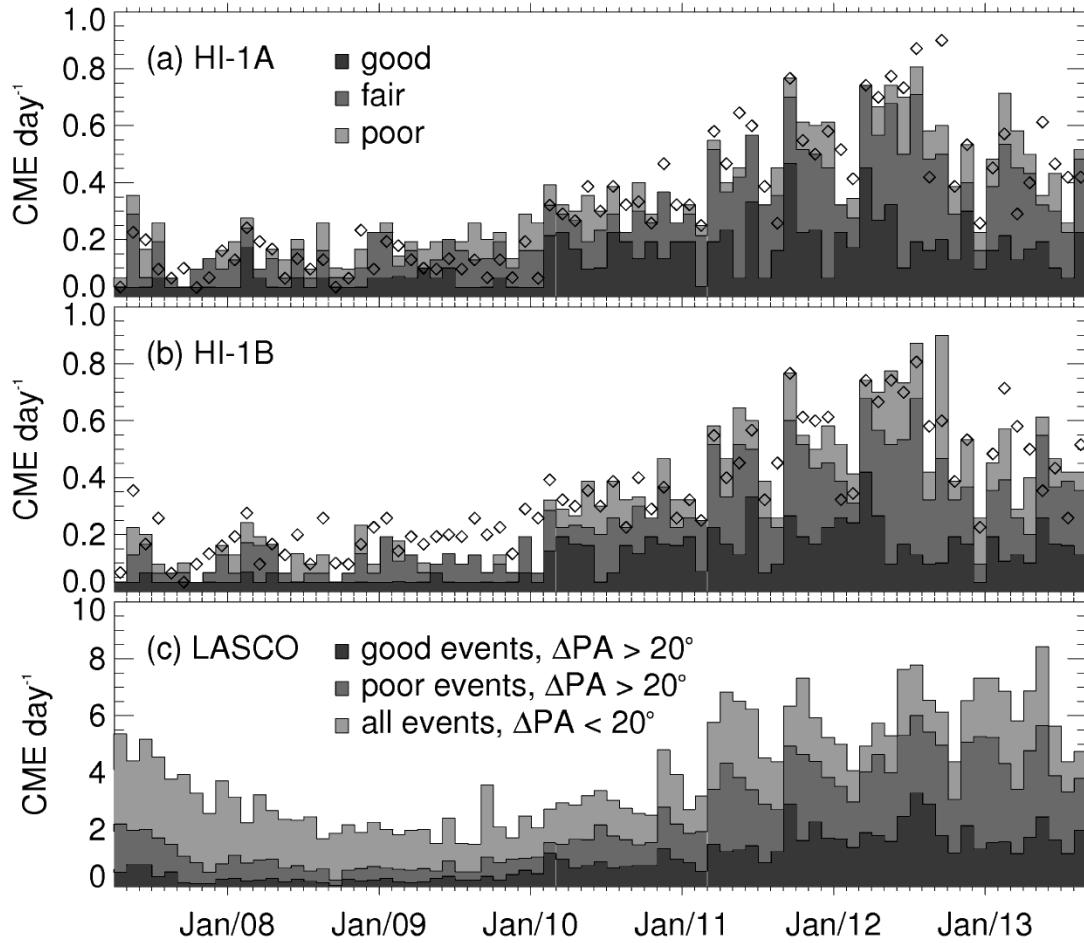
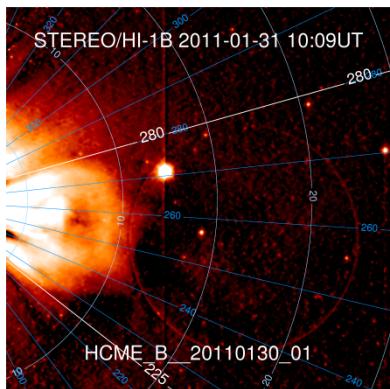
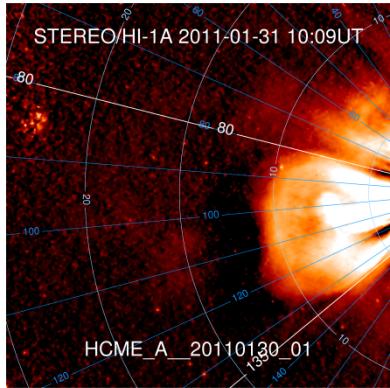
Contents

- A description of the basic HI CME catalogue
- Modelling techniques used to estimate CME kinematic properties from HI observations
- Statistical properties from the new CME catalogue
- Stereoscopic modelling of CMEs observed by both spacecraft
- Summary



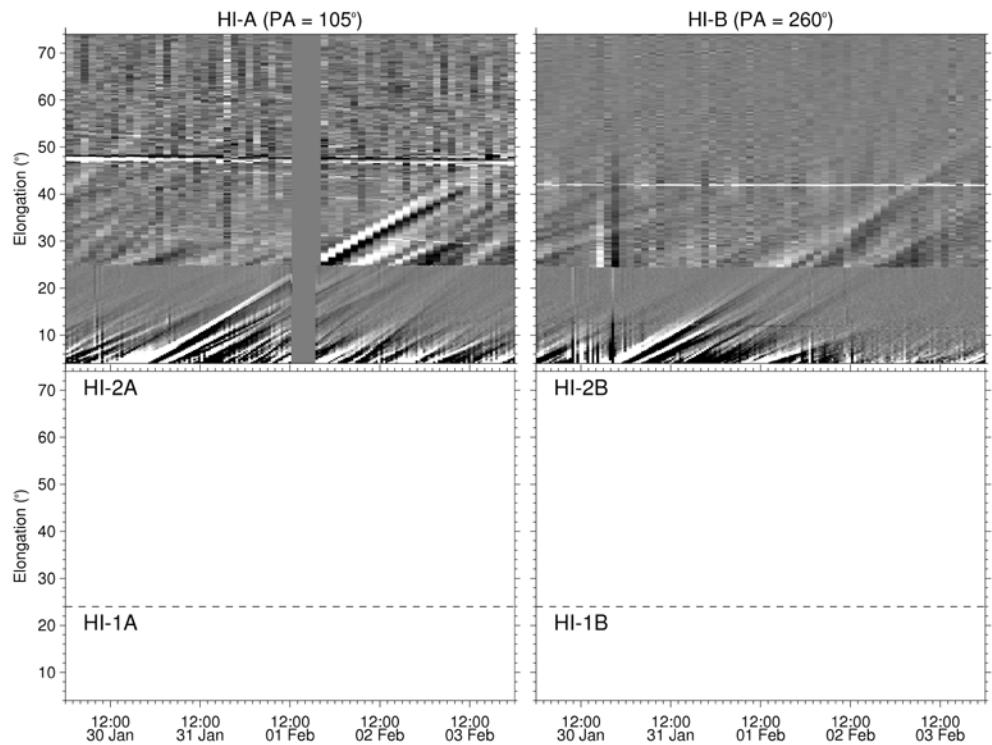
The HELCATS HI-Catalogue

Contains the basic observational properties of CMEs observed during the science phase of the STEREO mission (April 2007 - September 2014)



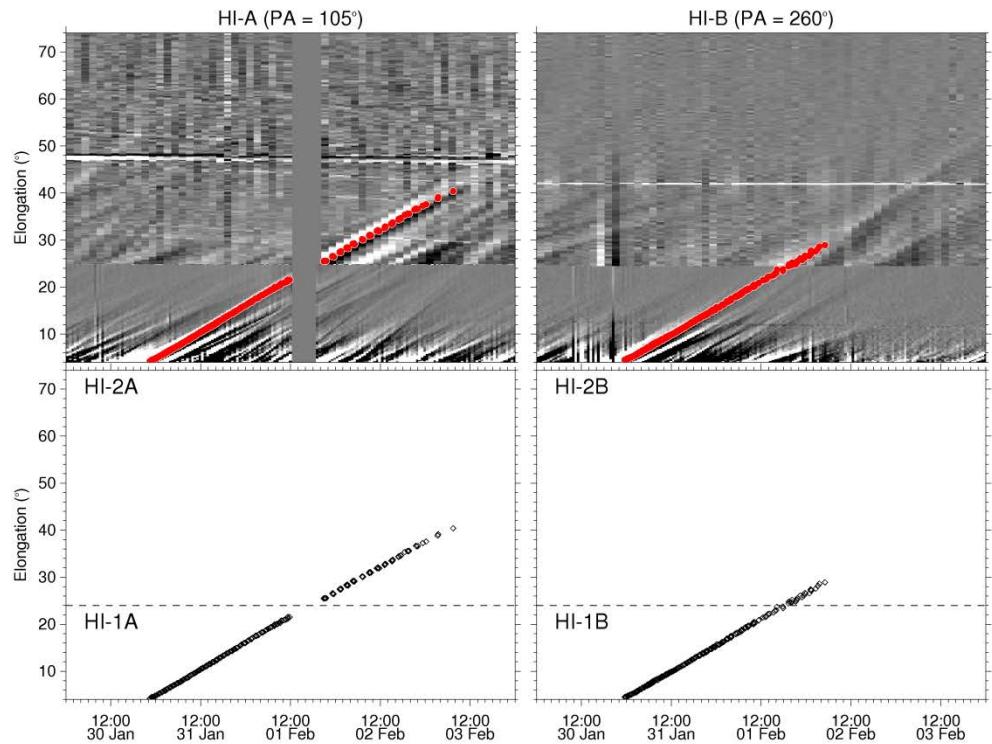
CME Tracking

- CMEs are identified in time/elongation plot



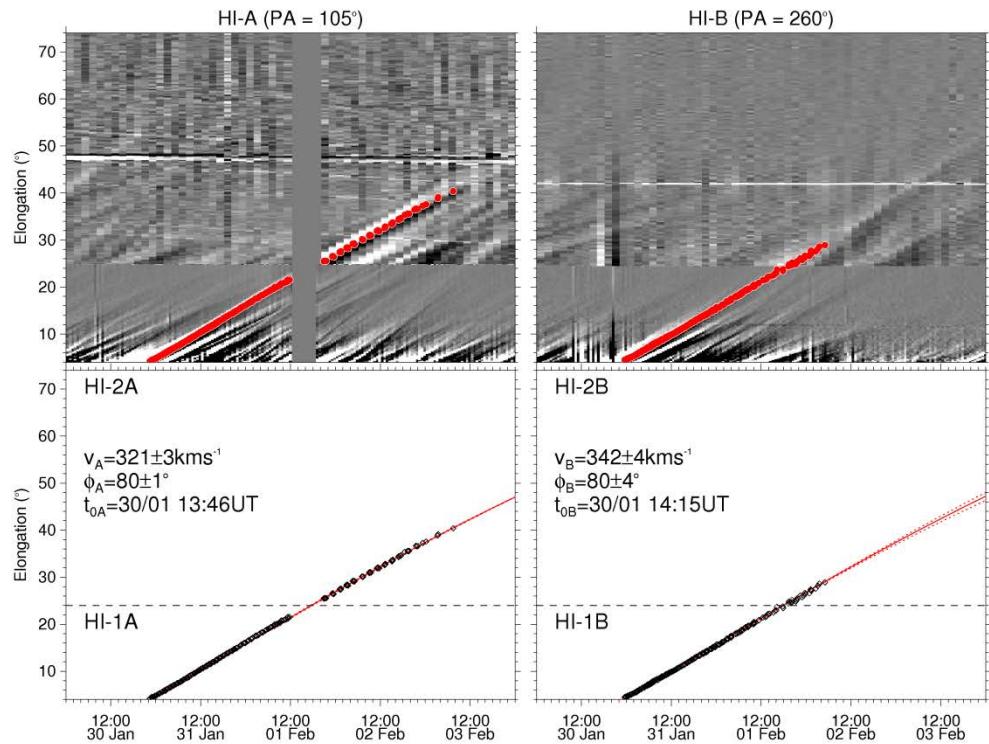
CME Tracking

- CMEs are identified in time/elongation plot
- Tracks are manually identified

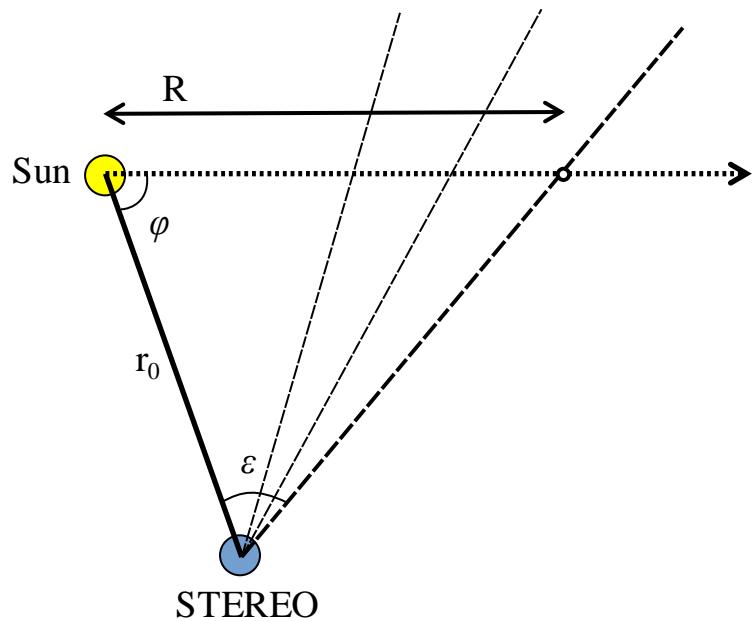


CME Tracking

- CMEs are identified in time/elongation plot
- Tracks are manually identified
- Kinematic properties are determined based on assumptions of CME morphology (Davies et al. 2012)

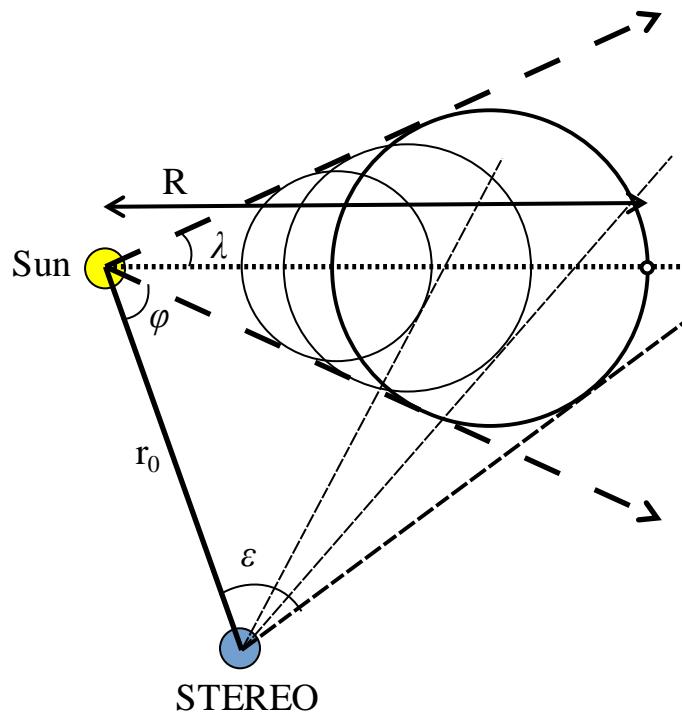


1st Model: Fixed-phi



$$v(t - t_0) = \frac{r_0 \sin(\varepsilon(t))}{\sin(\varepsilon(t) + \varphi)}$$

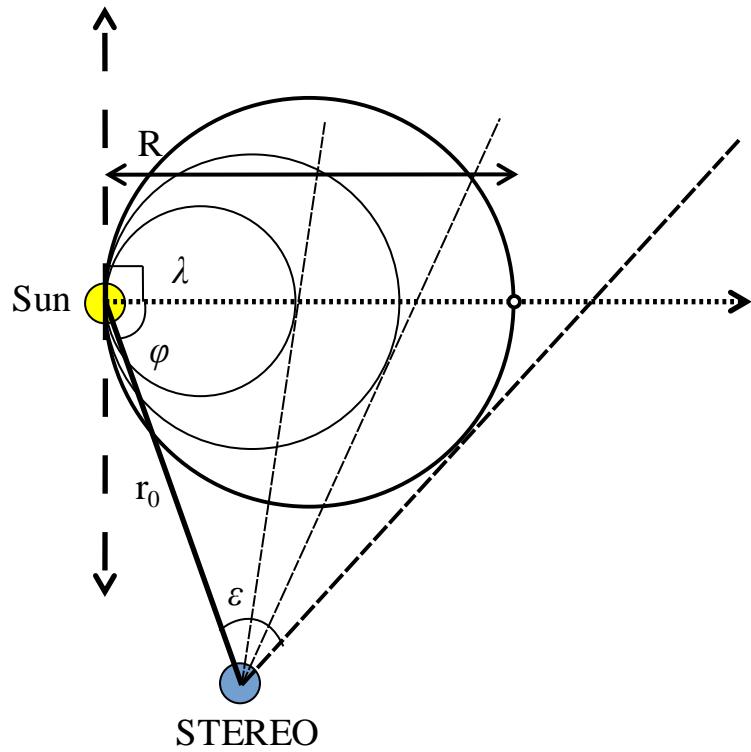
2nd Model: Self-Similar Expansion



$$v(t - t_0) = \frac{3r_0 \sin(\varepsilon(t))}{2\sin(\varepsilon(t) + \varphi) + \frac{1}{2}}$$

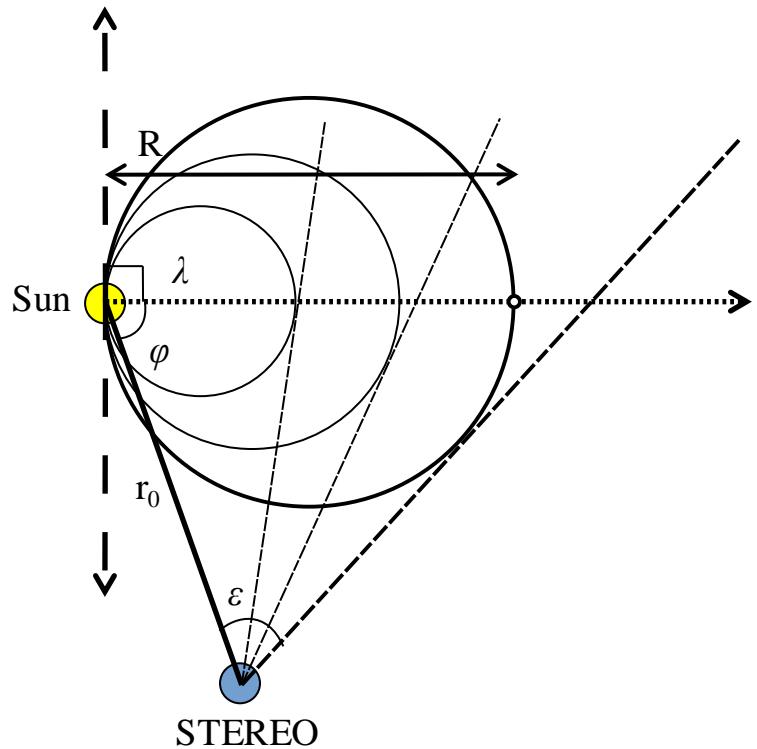


3rd Model: Harmonic Mean



$$v(t - t_0) = \frac{2r_0 \sin(\varepsilon(t))}{\sin(\varepsilon(t) + \varphi) + 1}$$

3rd Model: Harmonic Mean



$$v(t - t_0) = \frac{2r_0 \sin(\varepsilon(t))}{\sin(\varepsilon(t) + \varphi) + 1}$$

$$v(t - t_0) = \frac{r_0 \sin(\varepsilon(t))(1 + \sin(\lambda))}{\sin(\varepsilon(t) + \varphi) + \sin(\lambda)}$$

- | | |
|------------------------|------------------------|
| $\lambda = 0^\circ$; | fixed-phi |
| $\lambda = 30^\circ$; | self-similar expansion |
| $\lambda = 90^\circ$; | harmonic mean |

Catalogue of CME Kinematic Properties

Catalogue is complete for 1353 CMEs (Apr '07 - Sep '14)

SSE launch date range

SSE speed: 50 to 3600 kms⁻¹

SSE HEEQ Lon: -180 to 180 degrees

SSE HEEQ Lat: -45 to 45 degrees

From to



Show entries

Search: Show / hide columns

ID	SC	Quality	PA-fit	SSE speed [kms ⁻¹]	SSE Phi [deg]	SSE HEEQ Long [deg]	SSE HEEQ Lat [deg]	SSE Carr Long [deg]	SSE Launch [UTC]
HCME_A_20110122_01	A	fair	70	331	73	15	17	357	2011-01-21 18:47
HCME_A_20110124_01	A	good	95	379	72	13	-6	323	2011-01-24 03:39
HCME_A_20110125_01	A	good	60	362	80	31	23	323	2011-01-25 12:11
HCME_A_20110130_01	A	good	105	321	81	3	-15	229	2011-01-30 13:46
HCME_A_20110214_01	A	fair	100	299	77	13	-10	54	2011-02-14 15:30
HCME_A_20110214_02	A	good	90	417	70	16	0	42	2011-02-14 18:26
HCME_A_20110216_01	A	fair	70	750	144	-60	19	306	2011-02-16 04:25
HCME_A_20110217_01	A	fair	50	515	34	59	19	56	2011-02-16 23:26
HCME_A_20110225_01	A	fair	55	463	75	15	33	263	2011-02-25 06:46
HCME_A_20110226_01	A	fair	70	388	47	41	13	281	2011-02-25 20:17

Showing 231 to 240 of 1,315 entries (filtered from 1,330 total entries)

Select Save Print Previous 1 .. 23 24 25 .. 132 Next

speed

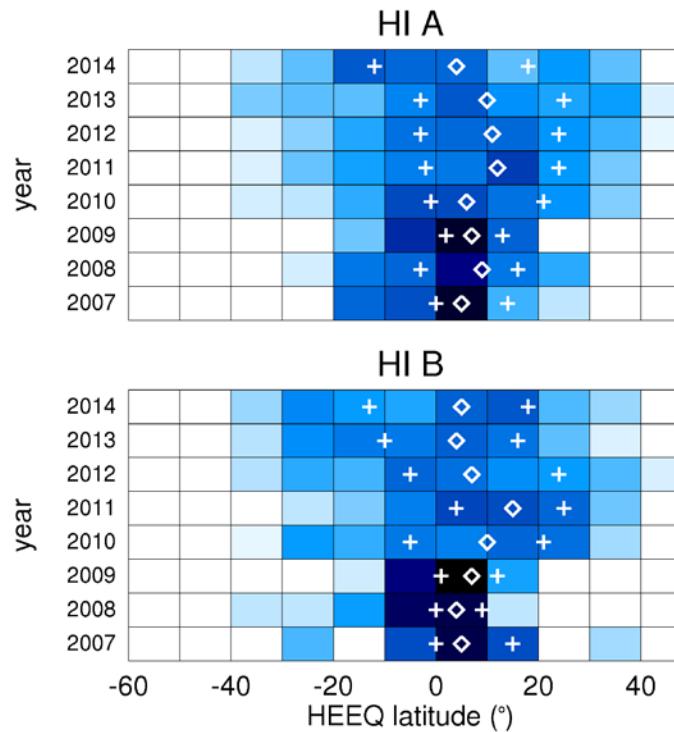
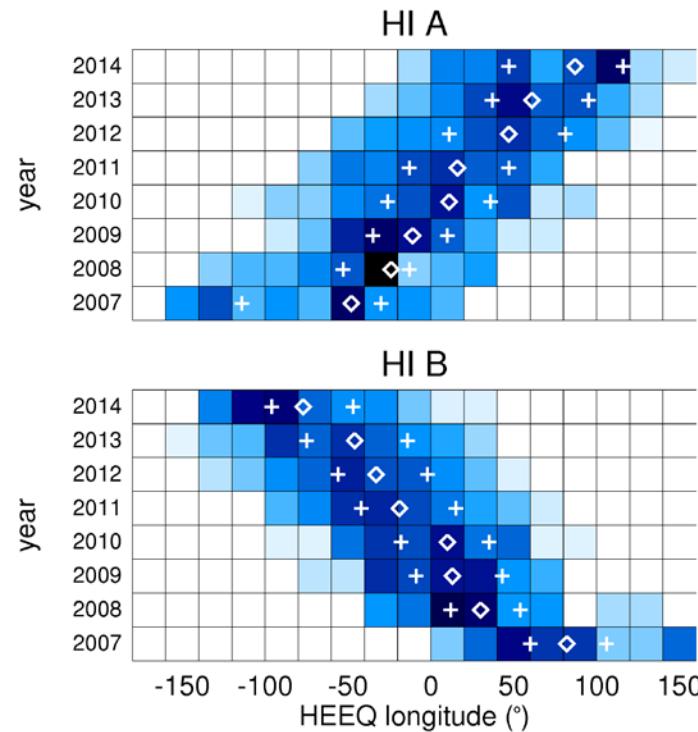
direction

launch time

http://www.helcats-fp7.eu/catalogues/wp3_cat.html



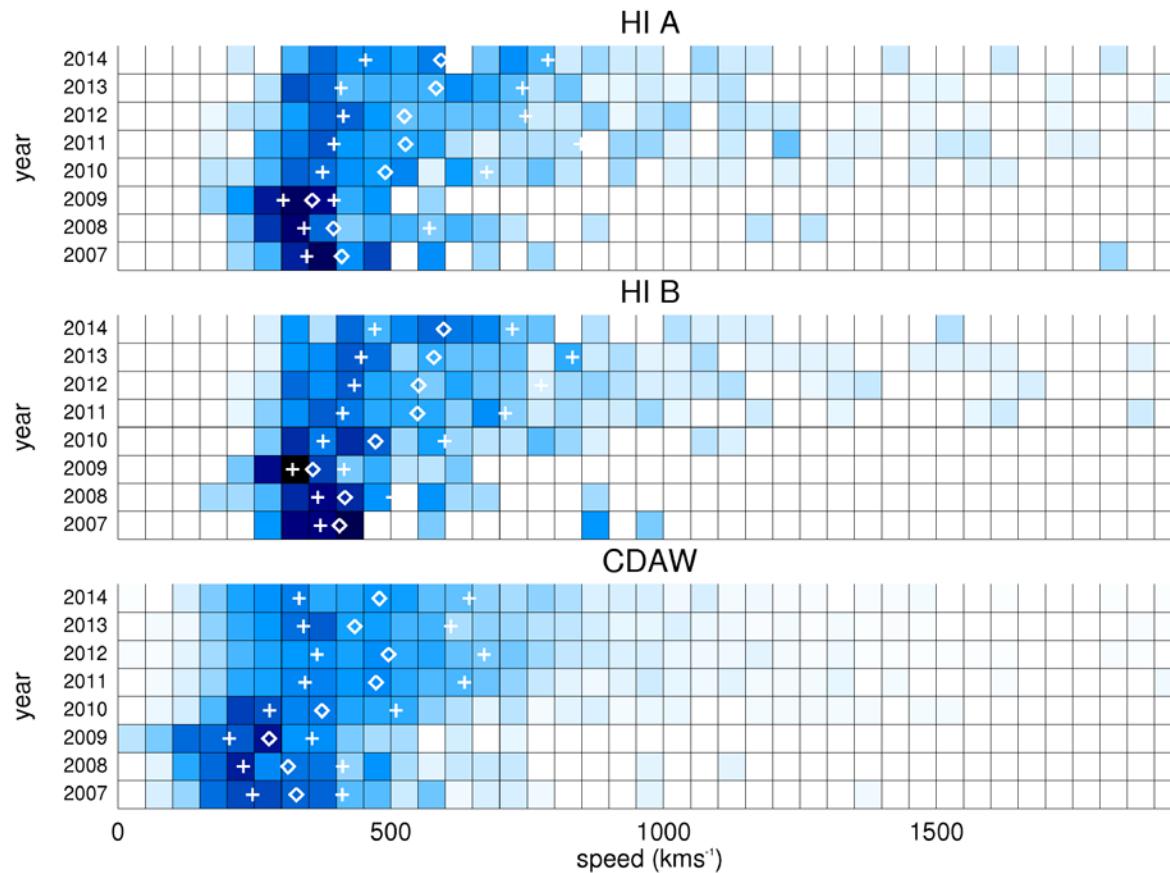
CME Statistical Properties



- CME latitude distributions consistent with established behaviour



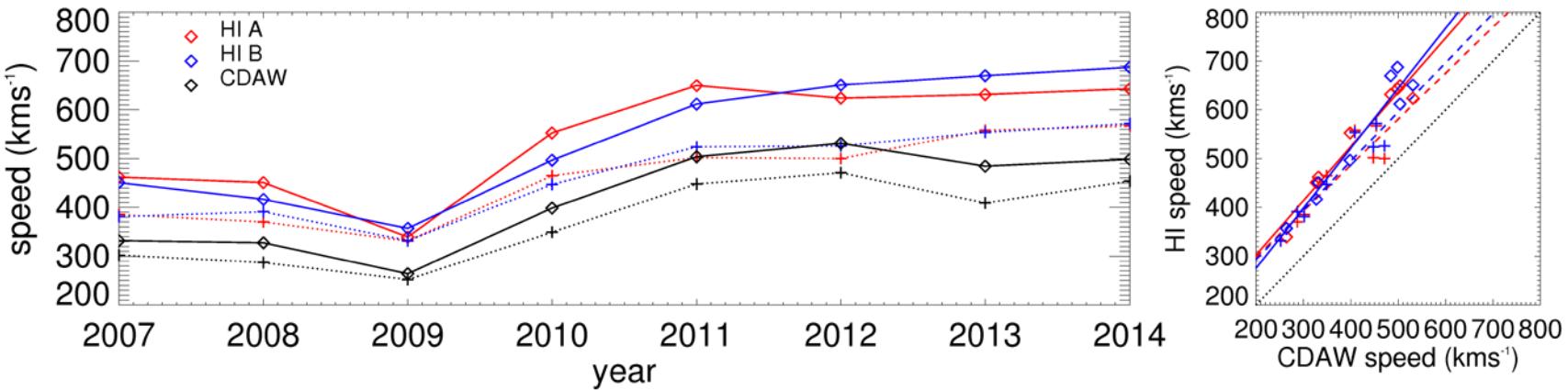
CME Statistical Properties



- Stacked annual speed histograms from HELCATS HI catalogues and LASCO CDAW catalogue (Yashiro et al. 2004)



CME Statistical Properties



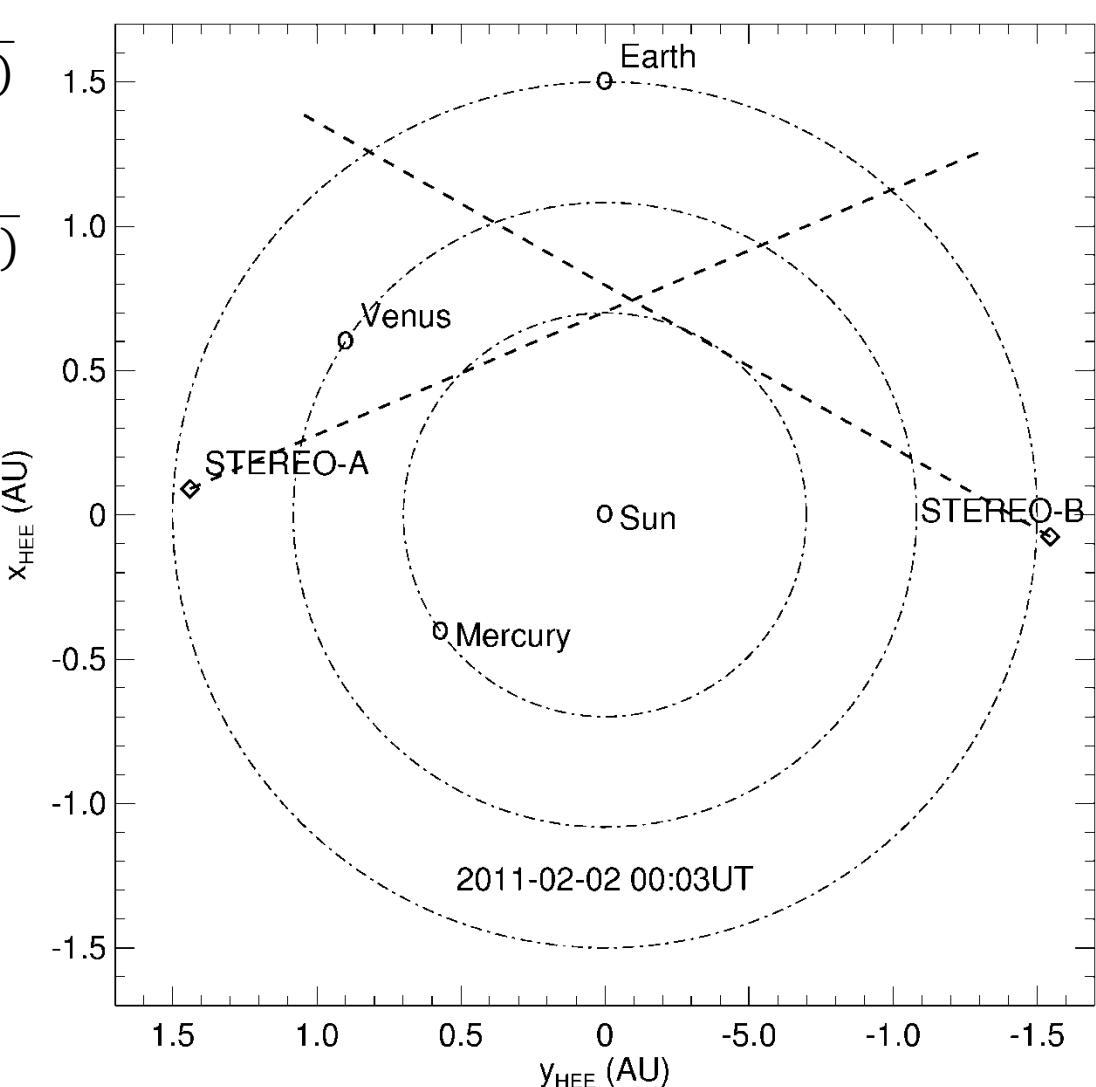
- Same solar cycle trend seen in HI and LASCO
- Mean speed are greater in HI by $\sim 120 \text{ km s}^{-1}$
- Due to ‘projection effects’ in coronagraph speed estimates

Stereoscopic Modelling

$$v(t - t_0) = \frac{r_A \sin(\varepsilon_A(t)) \sin(1 + \lambda)}{\sin(\varepsilon_A(t) + \varphi_A) + \sin(\lambda)}$$

$$= \frac{r_B \sin(\varepsilon_B(t)) \sin(1 + \lambda)}{\sin(\varepsilon_B(t) + \varphi_B) + \sin(\lambda)}$$

(Davies et al. 2013)

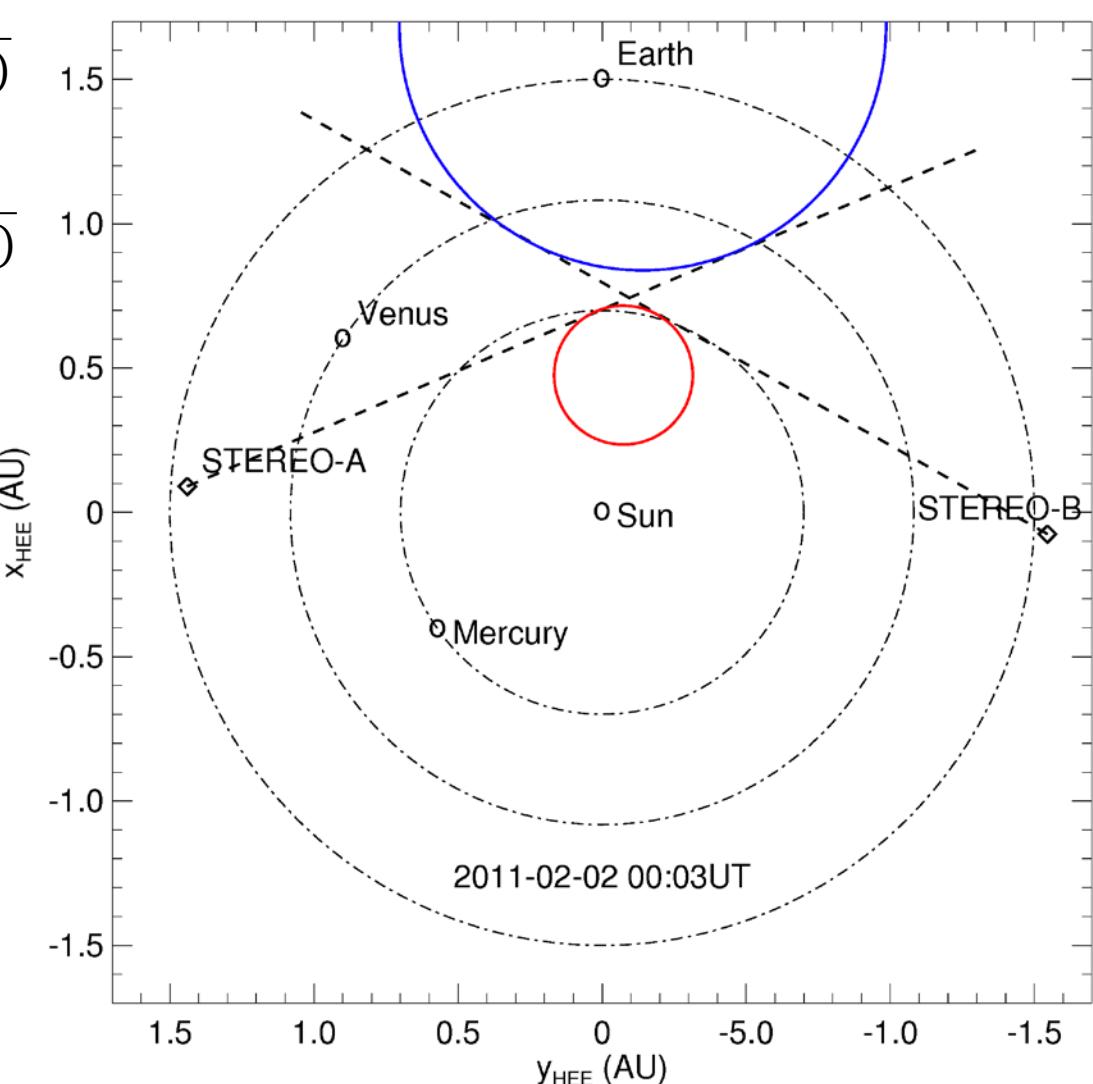


Stereoscopic Modelling

$$v(t - t_0) = \frac{r_A \sin(\varepsilon_A(t)) \sin(1 + \lambda)}{\sin(\varepsilon_A(t) + \varphi_A) + \sin(\lambda)}$$

$$= \frac{r_B \sin(\varepsilon_B(t)) \sin(1 + \lambda)}{\sin(\varepsilon_B(t) + \varphi_B) + \sin(\lambda)}$$

(Davies et al. 2013)

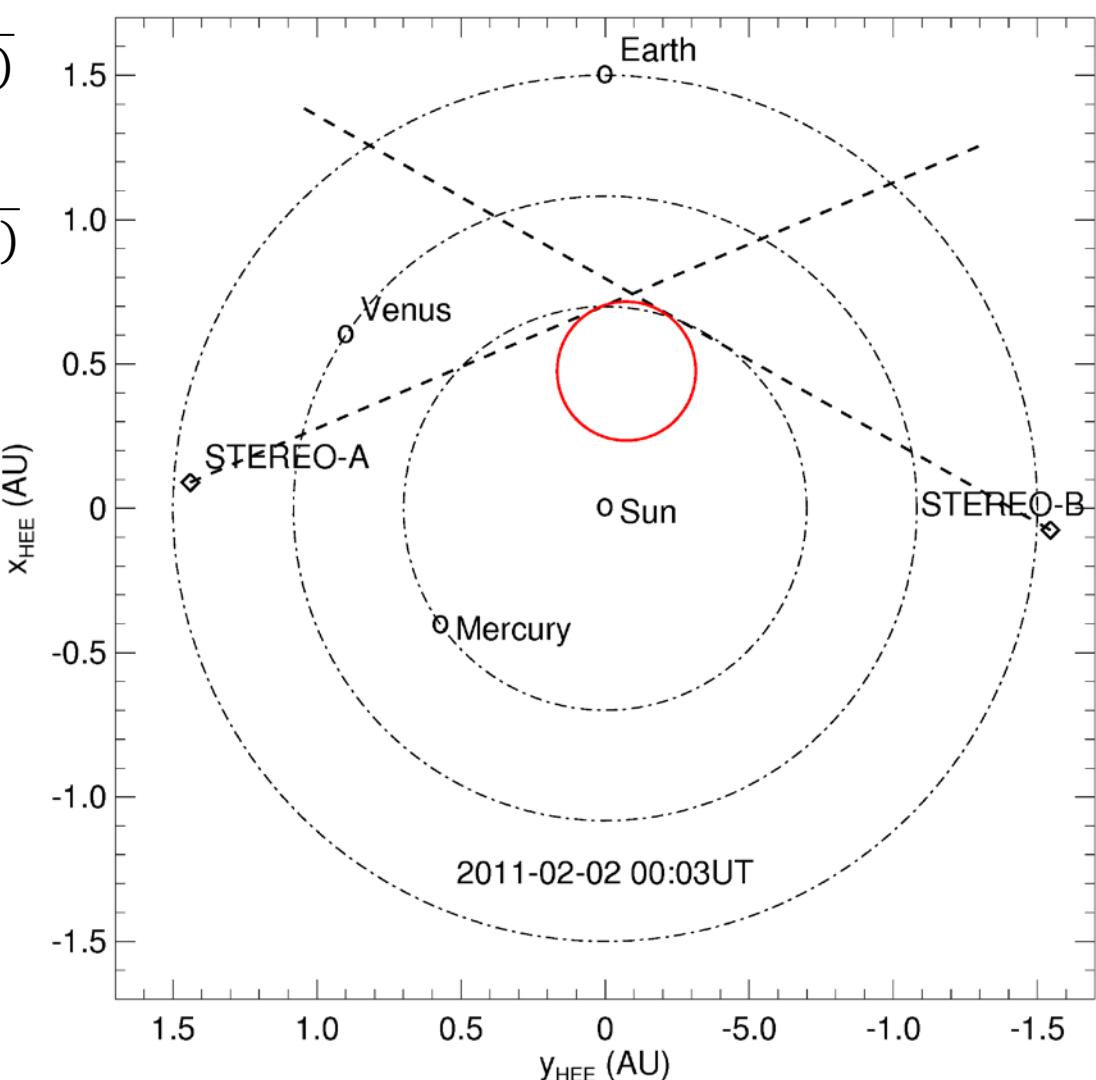


Stereoscopic Modelling

$$v(t - t_0) = \frac{r_A \sin(\varepsilon_A(t)) \sin(1 + \lambda)}{\sin(\varepsilon_A(t) + \varphi_A) + \sin(\lambda)}$$

$$= \frac{r_B \sin(\varepsilon_B(t)) \sin(1 + \lambda)}{\sin(\varepsilon_B(t) + \varphi_B) + \sin(\lambda)}$$

(Davies et al. 2013)

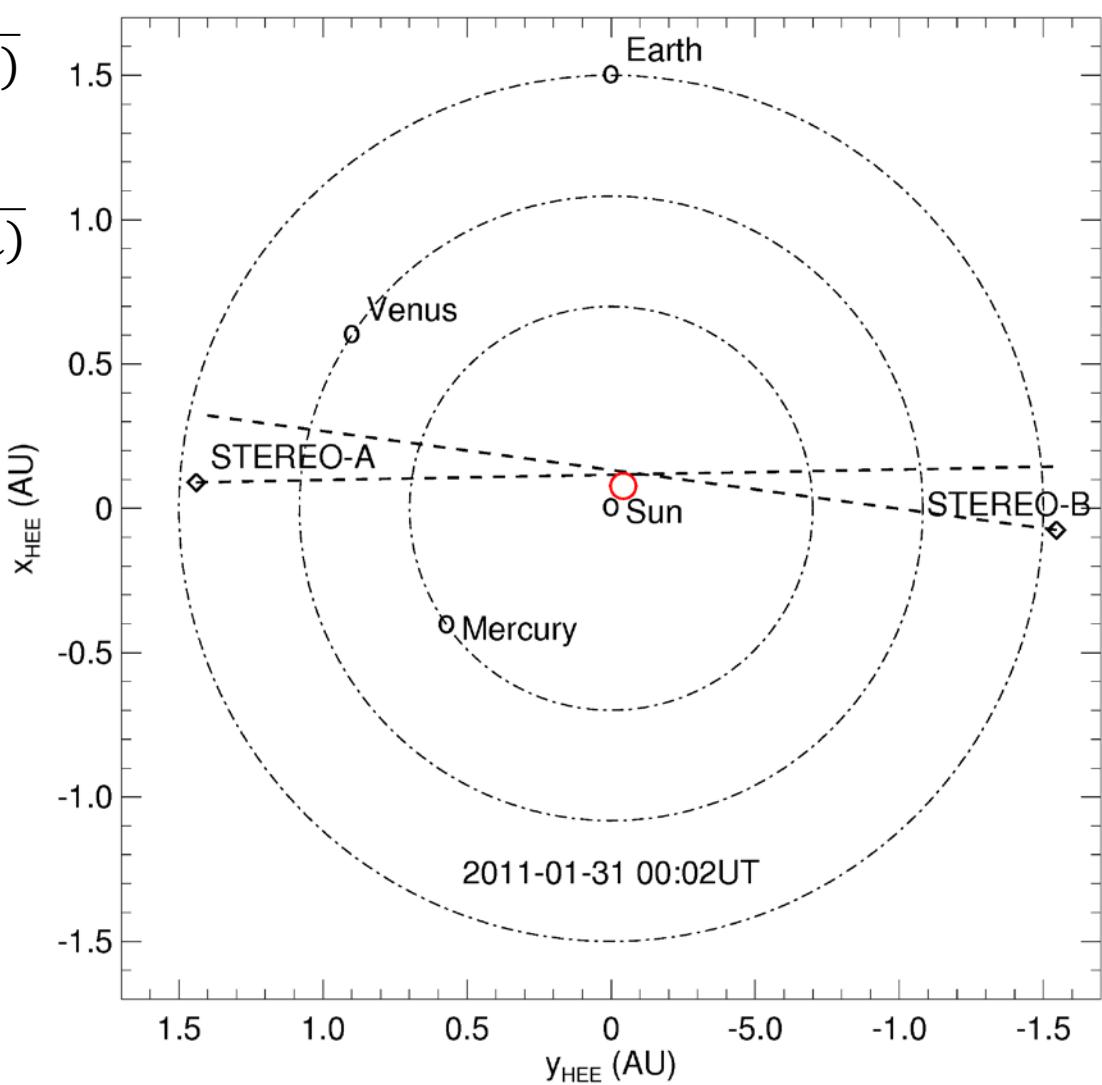


Stereoscopic Modelling

$$v(t - t_0) = \frac{r_A \sin(\varepsilon_A(t)) \sin(1 + \lambda)}{\sin(\varepsilon_A(t) + \varphi_A) + \sin(\lambda)}$$

$$= \frac{r_B \sin(\varepsilon_B(t)) \sin(1 + \lambda)}{\sin(\varepsilon_B(t) + \varphi_B) + \sin(\lambda)}$$

(Davies et al. 2013)

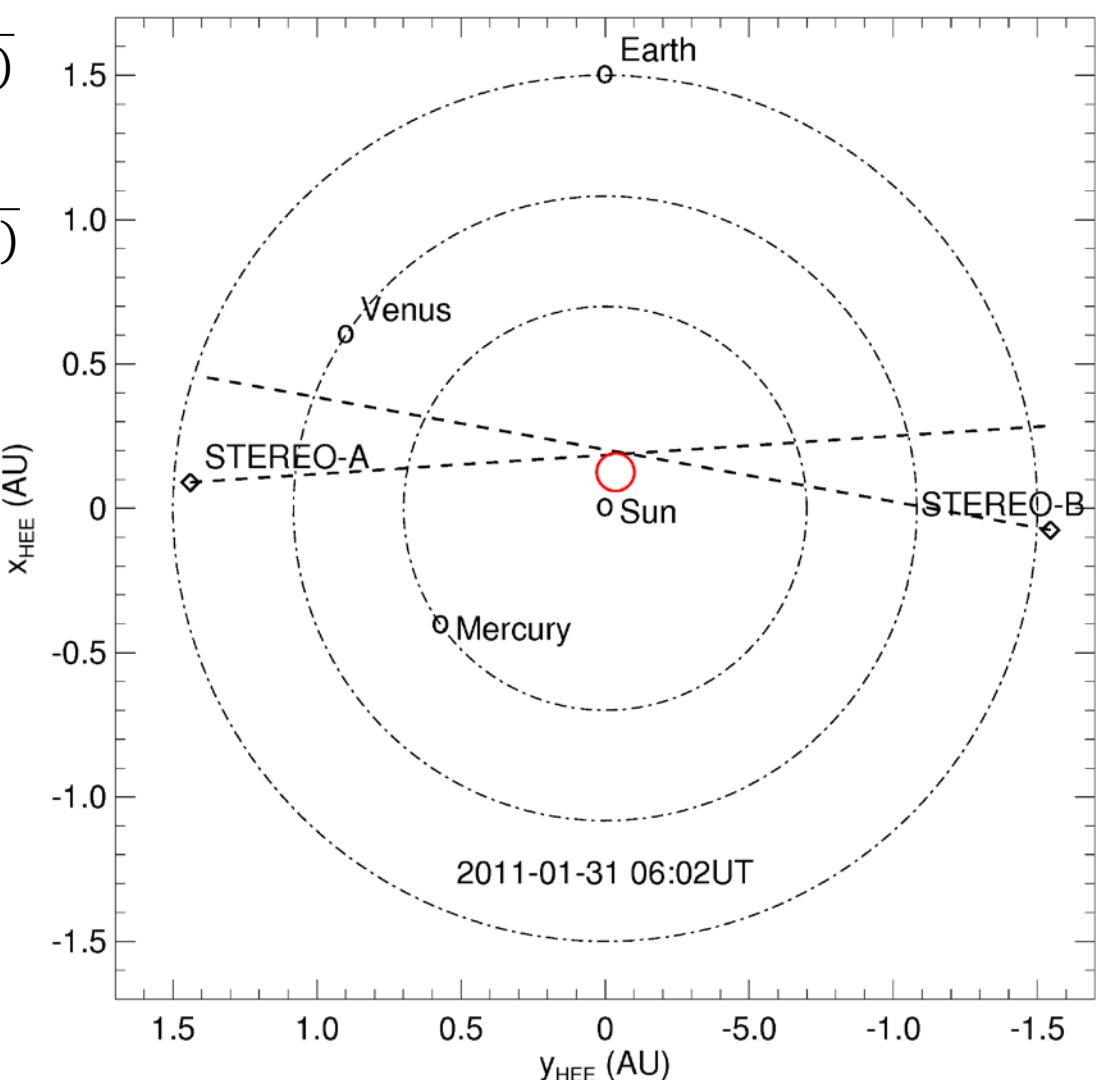


Stereoscopic Modelling

$$v(t - t_0) = \frac{r_A \sin(\varepsilon_A(t)) \sin(1 + \lambda)}{\sin(\varepsilon_A(t) + \varphi_A) + \sin(\lambda)}$$

$$= \frac{r_B \sin(\varepsilon_B(t)) \sin(1 + \lambda)}{\sin(\varepsilon_B(t) + \varphi_B) + \sin(\lambda)}$$

(Davies et al. 2013)

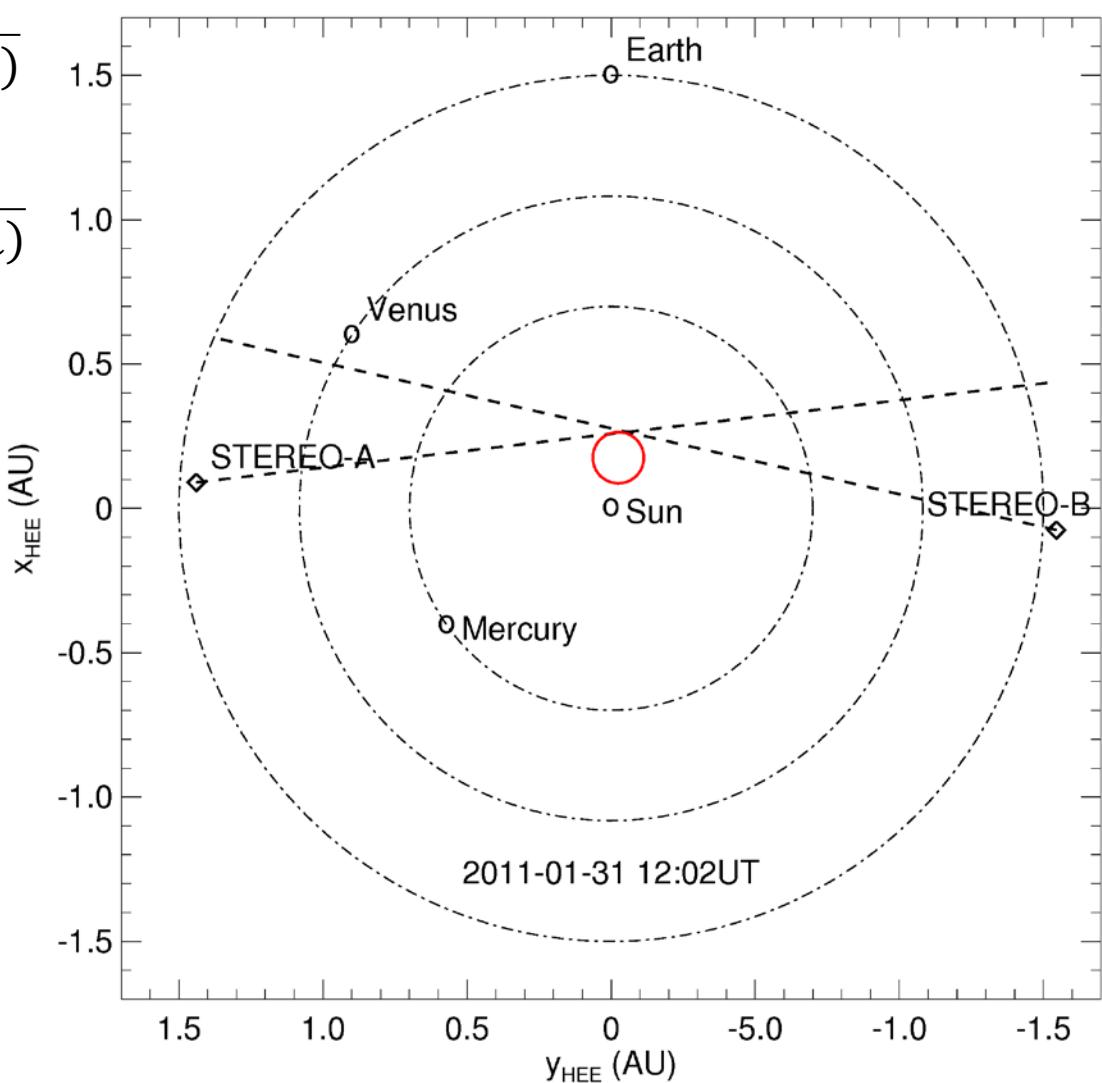


Stereoscopic Modelling

$$v(t - t_0) = \frac{r_A \sin(\varepsilon_A(t)) \sin(1 + \lambda)}{\sin(\varepsilon_A(t) + \varphi_A) + \sin(\lambda)}$$

$$= \frac{r_B \sin(\varepsilon_B(t)) \sin(1 + \lambda)}{\sin(\varepsilon_B(t) + \varphi_B) + \sin(\lambda)}$$

(Davies et al. 2013)

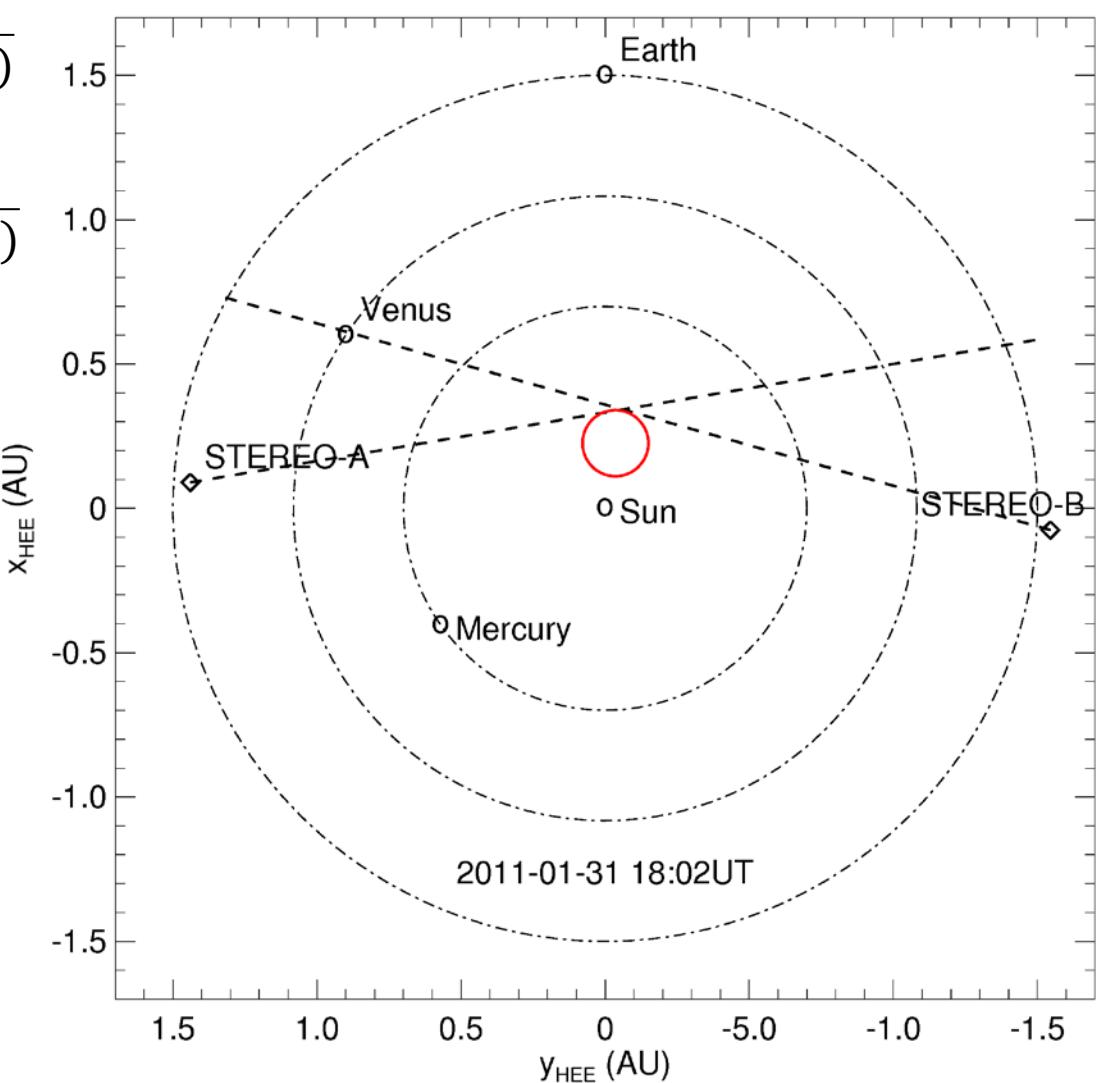


Stereoscopic Modelling

$$v(t - t_0) = \frac{r_A \sin(\varepsilon_A(t)) \sin(1 + \lambda)}{\sin(\varepsilon_A(t) + \varphi_A) + \sin(\lambda)}$$

$$= \frac{r_B \sin(\varepsilon_B(t)) \sin(1 + \lambda)}{\sin(\varepsilon_B(t) + \varphi_B) + \sin(\lambda)}$$

(Davies et al. 2013)

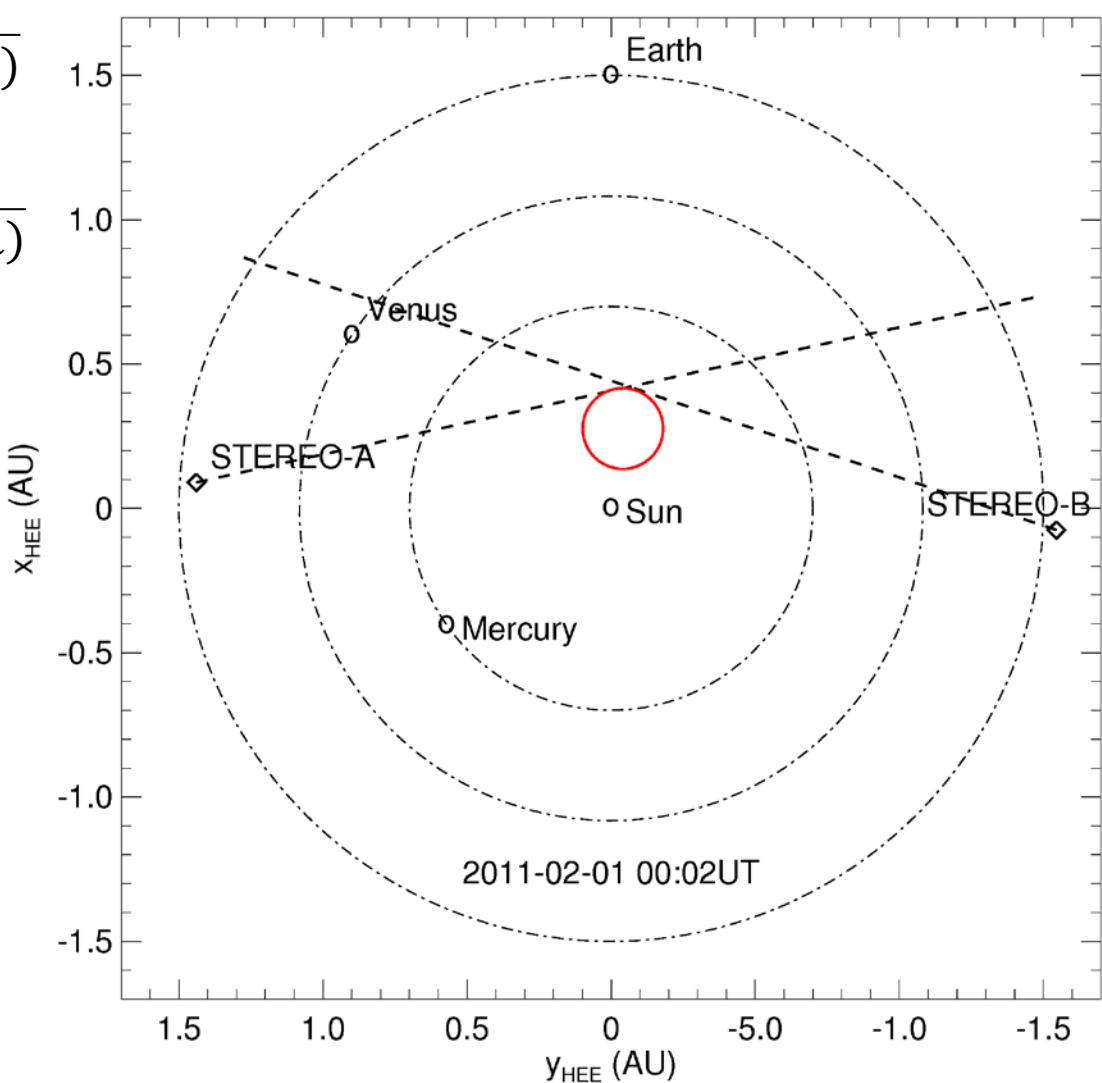


Stereoscopic Modelling

$$v(t - t_0) = \frac{r_A \sin(\varepsilon_A(t)) \sin(1 + \lambda)}{\sin(\varepsilon_A(t) + \varphi_A) + \sin(\lambda)}$$

$$= \frac{r_B \sin(\varepsilon_B(t)) \sin(1 + \lambda)}{\sin(\varepsilon_B(t) + \varphi_B) + \sin(\lambda)}$$

(Davies et al. 2013)

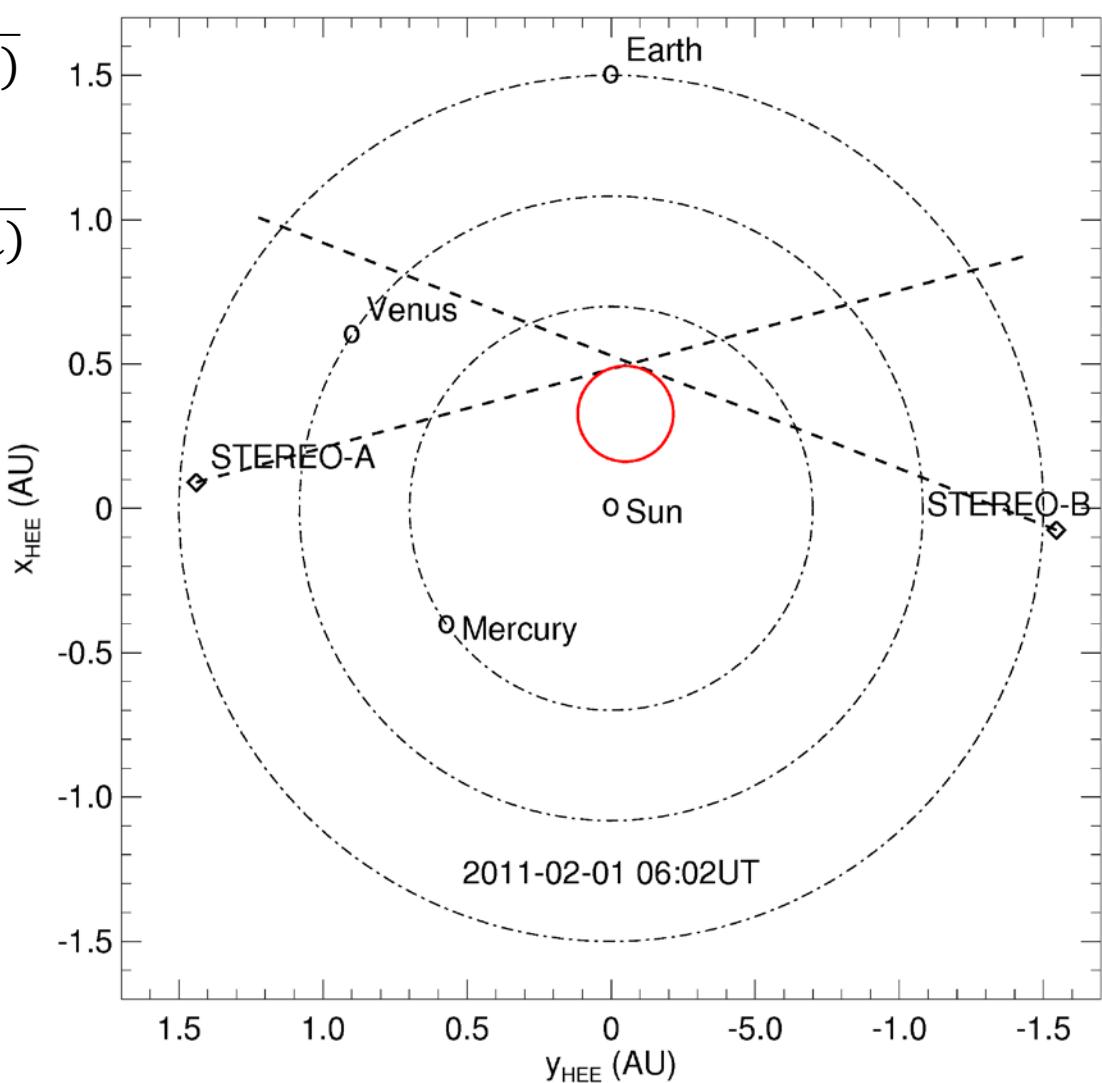


Stereoscopic Modelling

$$v(t - t_0) = \frac{r_A \sin(\varepsilon_A(t)) \sin(1 + \lambda)}{\sin(\varepsilon_A(t) + \varphi_A) + \sin(\lambda)}$$

$$= \frac{r_B \sin(\varepsilon_B(t)) \sin(1 + \lambda)}{\sin(\varepsilon_B(t) + \varphi_B) + \sin(\lambda)}$$

(Davies et al. 2013)

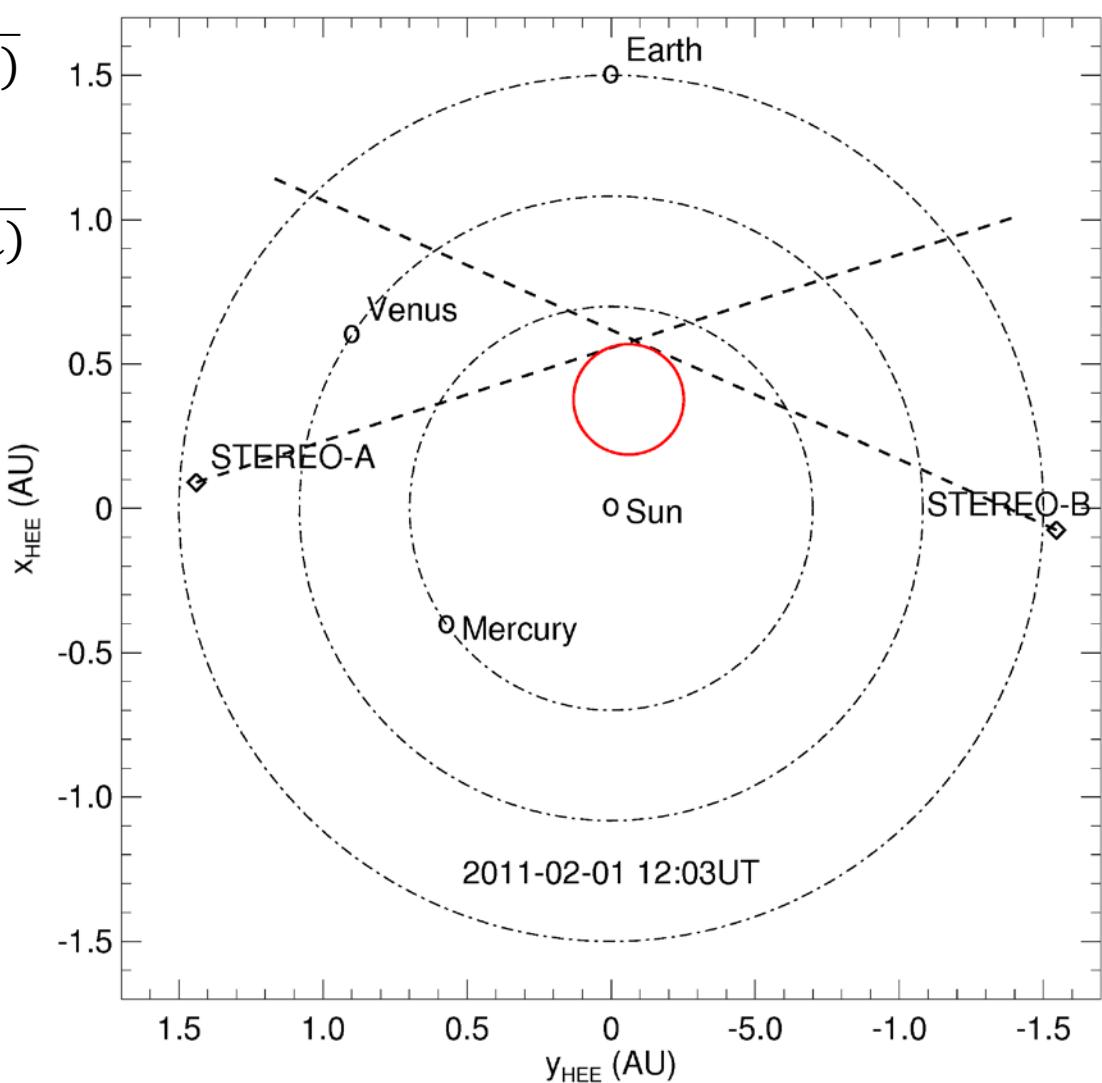


Stereoscopic Modelling

$$v(t - t_0) = \frac{r_A \sin(\varepsilon_A(t)) \sin(1 + \lambda)}{\sin(\varepsilon_A(t) + \varphi_A) + \sin(\lambda)}$$

$$= \frac{r_B \sin(\varepsilon_B(t)) \sin(1 + \lambda)}{\sin(\varepsilon_B(t) + \varphi_B) + \sin(\lambda)}$$

(Davies et al. 2013)

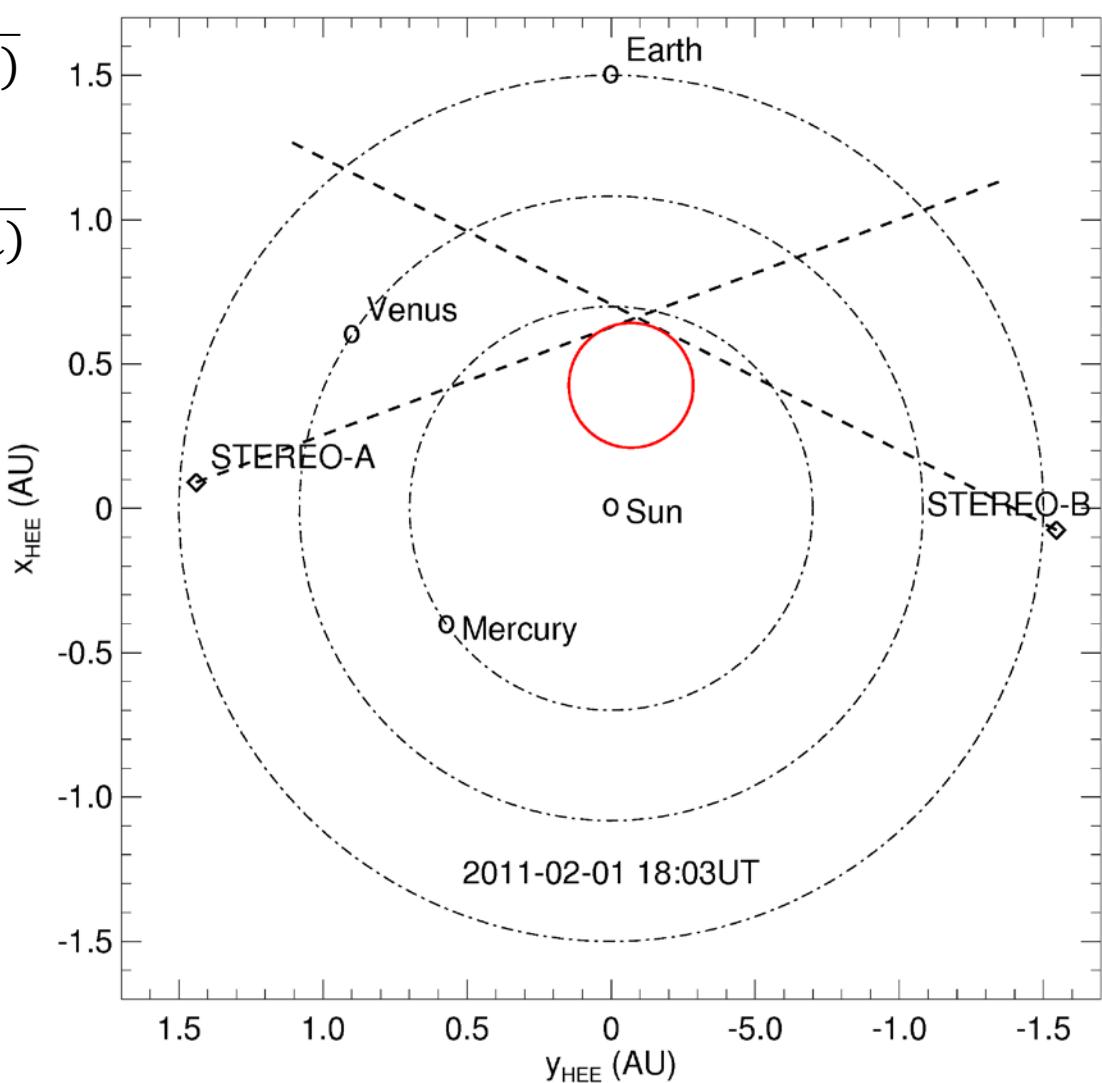


Stereoscopic Modelling

$$v(t - t_0) = \frac{r_A \sin(\varepsilon_A(t)) \sin(1 + \lambda)}{\sin(\varepsilon_A(t) + \varphi_A) + \sin(\lambda)}$$

$$= \frac{r_B \sin(\varepsilon_B(t)) \sin(1 + \lambda)}{\sin(\varepsilon_B(t) + \varphi_B) + \sin(\lambda)}$$

(Davies et al. 2013)

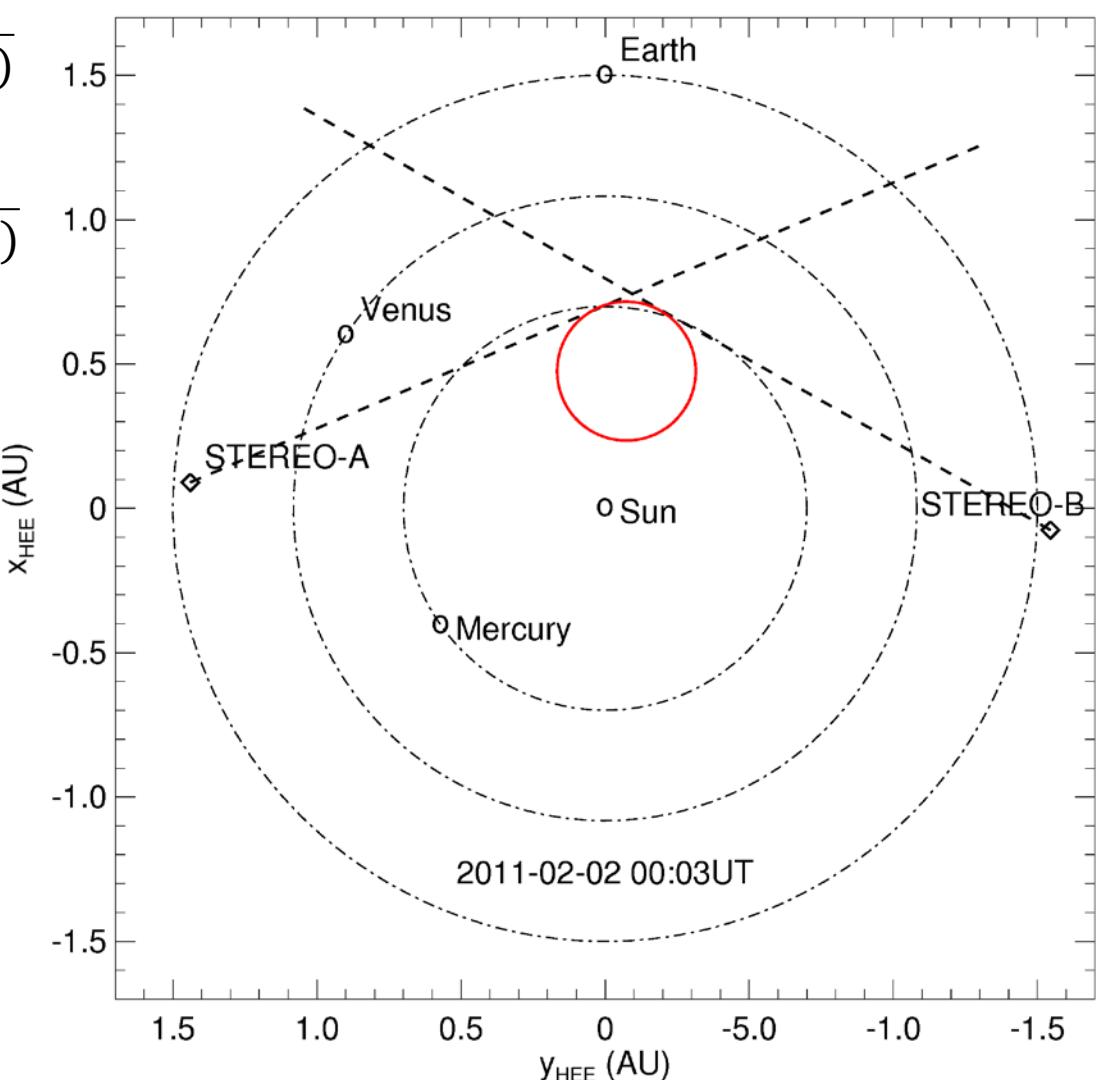


Stereoscopic Modelling

$$v(t - t_0) = \frac{r_A \sin(\varepsilon_A(t)) \sin(1 + \lambda)}{\sin(\varepsilon_A(t) + \varphi_A) + \sin(\lambda)}$$

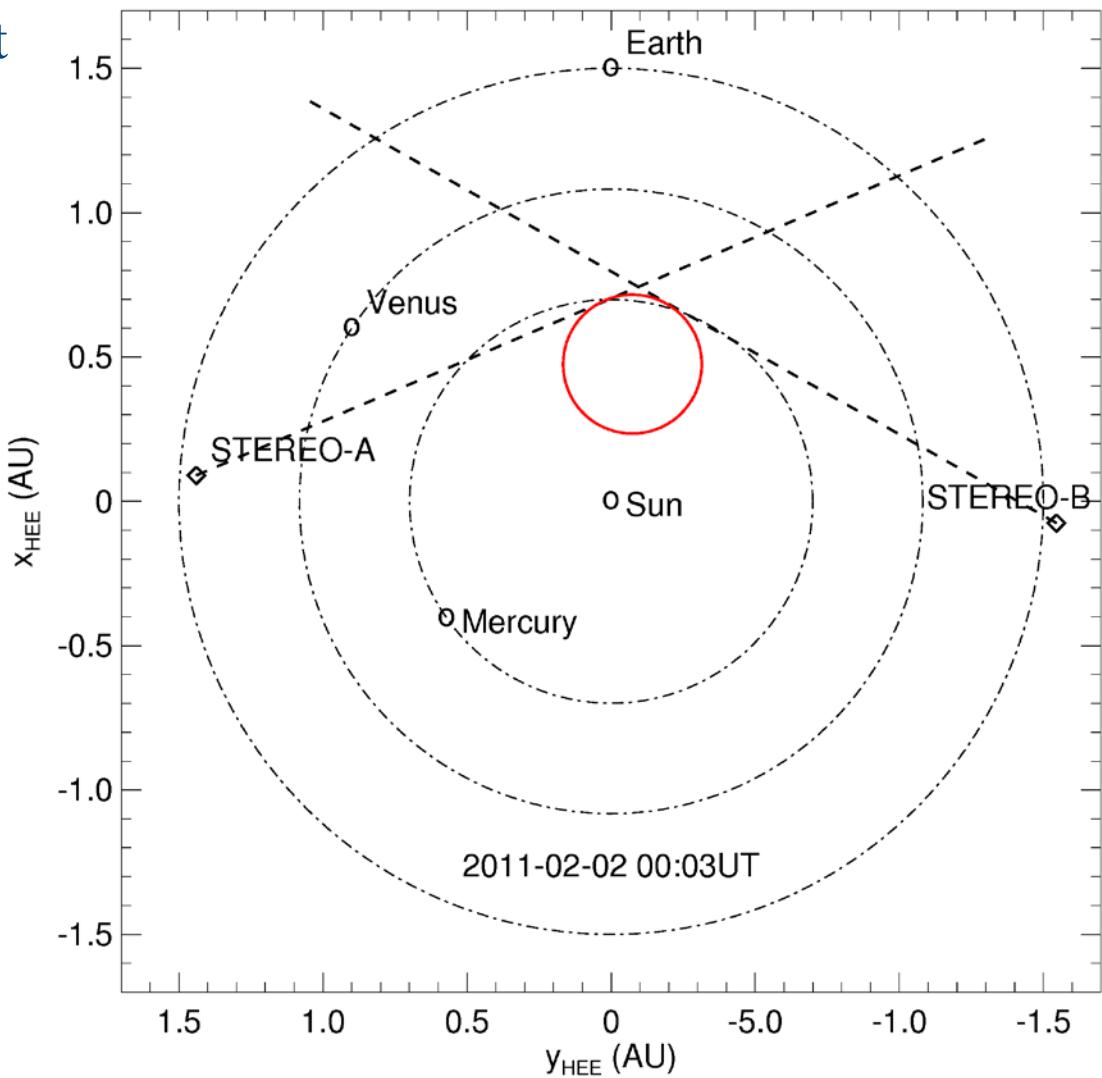
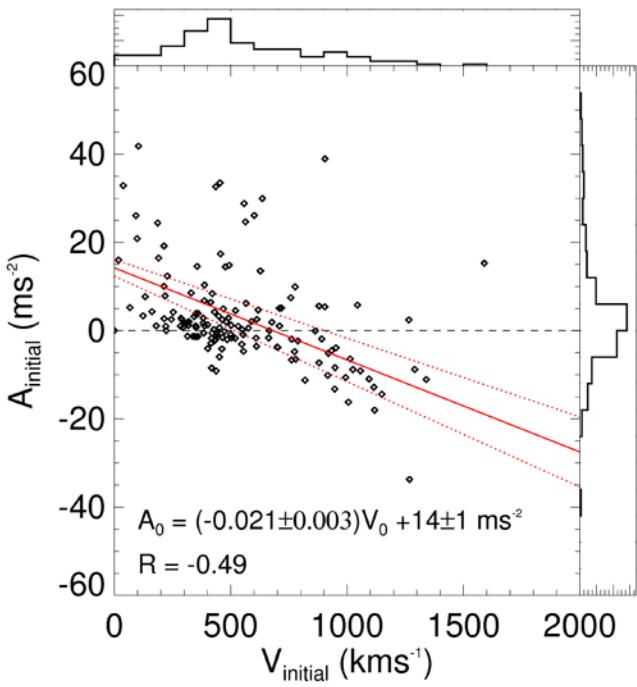
$$= \frac{r_B \sin(\varepsilon_B(t)) \sin(1 + \lambda)}{\sin(\varepsilon_B(t) + \varphi_B) + \sin(\lambda)}$$

(Davies et al. 2013)



Stereoscopic Modelling

- So far a preliminary subset of 393 CMEs exists



Summary

- Heliospheric Cataloguing Analysis and Techniques Service
 - 1901 CMEs identified throughout the lifetime of HI instruments
 - 1353 CMEs tracked and kinematic properties determined
 - Catalogues agree with established CME behaviour, but show greater speeds
 - Stereoscopic analysis allows study of CME accelerations/deflections in HI

Thanks for listening

